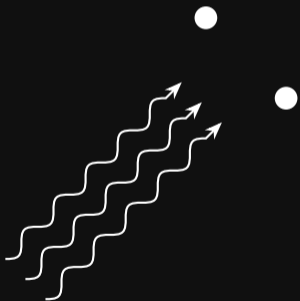


The Memory Effect and What it Means for Humankind

Níckolas de Aguiar Alves
Federal University of ABC



Don't Forget About Memory



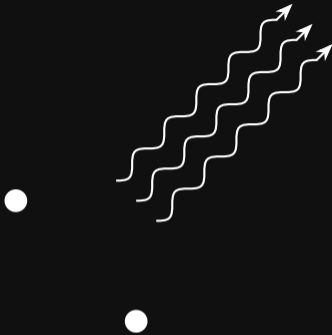
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Consider two inertial detectors “near infinity”. If a gravitational wave passes through them, what happens?



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In general, there will be a permanent displacement between the detectors. This is known as the (linear) memory effect.

How much deviation?

$$\frac{d^2 \Delta x^j}{dt^2} = -R_{0i0}{}^j d^i$$

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$$\frac{d^2 \Delta x^j}{dt^2} = -R_{0i0}{}^j d^i$$

We can calculate the deviation using the geodesic deviation equation. d^i is the initial deviation, Δx^j is the deviation as a function of time.

$$R_{0i0j} = -\frac{1}{r} \frac{d^4 Q_{ij}^{\text{TT}}}{dt^4} (t - r)$$

$$Q_{ij}(t) = \int T^{00}(t, \mathbf{x}) x_i x_j d^3 x$$

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$$R_{0i0j} = \frac{1}{r} \frac{d^4 Q_{ij}^{TT}}{dt^4} (t-r)$$

$$Q_{ij}(t) = \int T^{00}(t, \mathbf{x}) x_i x_j d^3x$$

At large r , the relevant Riemann tensor component can be computed from the transverse-traceless projection of the quadrupole moment. Notice we are ignoring non-radiative modes.

$$\frac{d^2 \Delta x^j}{dt^2} = \frac{1}{r} \frac{d^4 Q^{ijTT}}{dt^4} d_i$$

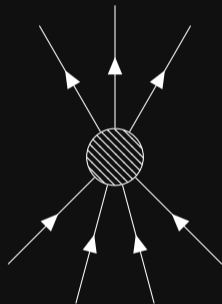
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$$\frac{d^2 \Delta x^i}{dt^2} = \frac{1}{r} \frac{d^4 Q^{ij\pi\pi}}{dt^4} d_i$$

The geodesic deviation equation in terms of the quadrupole moment.





For concreteness, consider the gravitational scattering of a number of bodies. We assume the bodies are asymptotically inertial *in the background Minkowski spacetime*.

$$\frac{d^2 Q^{ij}}{dt^2} = \sum_k m_k v_k^i v_k^j$$

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$$\frac{d^2 Q^{ij}}{dt^2} = \sum_k m_k v_k^i v_k^j$$

For this scattering setup, the second derivative of the quadrupole moment is given by a simple expression involving the masses and velocities of each body. At early and late times, the third derivative of the quadrupole moment vanishes.

$$\Delta x^j = \frac{d_j}{r} \frac{d^2 Q^{ij\text{TT}}}{dt^2} \Bigg|_{-\infty}^{\infty}$$

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$$\Delta x^j = \frac{d_i}{r} \frac{d^2 Q^{ijTT}}{dt^2} \Bigg|_{-\infty}^{\infty}$$

The generic memory deviation.

$$\Delta x^j = \frac{d_i}{r} \left[\sum_{k_f} m_{k_f} v_{k_f}^i v_{k_f}^j - \sum_{k_0} m_{k_0} v_{k_0}^i v_{k_0}^j \right]^{TT}$$

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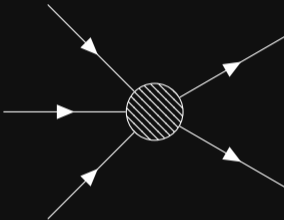
$$\Delta x^i = \frac{d_i}{r} \left[\sum_{R_f} m_{R_f} v_{R_f}^i v_{R_f}^j - \sum_{R_0} m_{R_0} v_{R_0}^i v_{R_0}^j \right]^{TT}$$

The Braginsky–Thorne formula for the memory due to a scattering process.

Note: there is also a nonlinear memory effect

Note: the nonlinear memory effect can be understood as the memory effect due to the propagation of gravitational waves themselves

And Now a Brief Message from Our Sponsor
Infrared Structure of Gravity

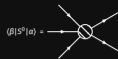
$$\langle \beta | S^0 | \alpha \rangle =$$
A Feynman diagram representing a scattering process. It features a central circle filled with diagonal hatching. Four lines with arrowheads point towards this central circle from the left, and four lines with arrowheads point away from it towards the right. The diagram is positioned to the right of an equals sign, which is part of an equation.

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Consider an arbitrary $\alpha \rightarrow \beta$ scattering in a theory that includes gravitons. Consider the amplitude for this scattering assuming no soft gravitons are exchanged or emitted.

$$\langle \beta | S | \alpha \rangle =$$

$$+ \dots$$

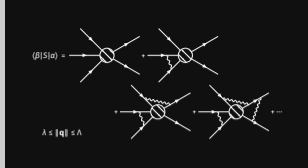
$\lambda \leq \| \mathbf{q} \| \leq \Lambda$

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Consider now the $\alpha \rightarrow \beta$ scattering with the exchange of an arbitrary number of soft gravitons (defined to have momentum $\lambda \leq \|\mathbf{q}\| \leq \Lambda$). Λ defines what is a soft graviton, while λ is an infrared cutoff to show the effect of the divergences.

$$|\langle \beta | S | \alpha \rangle|^2 = |\langle \beta | S^0 | \alpha \rangle|^2 \left(\frac{\lambda}{\Lambda} \right)^B \xrightarrow{\lambda \rightarrow 0} 0$$

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$$|\langle \beta | S | \alpha \rangle|^2 = |\langle \beta | S^0 | \alpha \rangle|^2 \left(\frac{\Lambda}{\lambda} \right)^0 \xrightarrow{\lambda \rightarrow 0} 0$$

The rate with soft graviton exchanges is simply related to the original transition rate. However, when removing the infrared cutoff, the theory is plagued with infrared divergences that force the rate to vanish.

$$\langle \beta | S^{\text{gr}} | \alpha \rangle =$$

$$+ \dots$$

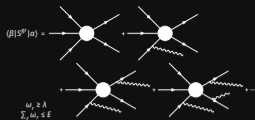
$\omega_r \geq \lambda$
 $\sum_r \omega_r \leq E$

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Consider now the $\alpha \rightarrow \beta$ scattering with the emission of an arbitrary number of soft gravitons (defined to have energies $\omega \geq \lambda$ and total summed energy $\sum_r \omega_r \leq E$). E restricts the total energy “lost” to soft gravitons, while λ is an infrared cutoff to show the effect of the divergences.

$$|\langle \beta | S^{\text{gr}} | \alpha \rangle|^2 = |\langle \beta | S | \alpha \rangle|^2 b(B) \left(\frac{E}{\lambda} \right)^B \xrightarrow{\lambda \rightarrow 0} \infty$$

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$$|\langle \beta | S^{\alpha} | \alpha \rangle|^2 = |\langle \beta | S | \alpha \rangle|^2 b(B) \left(\frac{E}{\lambda} \right)^B \xrightarrow{\lambda \rightarrow 0} \infty$$

The rate with soft graviton emissions is simply related to the original transition rate. However, when removing the infrared cutoff, the theory is plagued with infrared divergences that lead the rate to a divergence.

$$|\langle \beta | S^{\text{gr}} | \alpha \rangle|^2 = |\langle \beta | S^0 | \alpha \rangle|^2 b(B) \left(\frac{E}{\Lambda} \right)^B$$

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$$|\langle \beta | S^{\mathcal{B}} | \alpha \rangle|^2 = |\langle \beta | S^0 | \alpha \rangle|^2 b(\mathcal{B}) \left(\frac{E}{\Lambda} \right)^{\mathcal{B}}$$

Combining the virtual and real calculations, the infrared divergences cancel out and the inclusive result is finite.

How are these calculations performed?

The dominance of the $1/(p \cdot q)$ pole in (2.5) implies that the effects of attaching one soft-graviton line to an arbitrary diagram is to supply a factor equal to the sum of (2.5) over all external lines in the diagram

$$(8\pi G)^{1/2} \sum_n \eta_n p_n^\mu p_n^\nu / [p_n \cdot q - i\eta_n \epsilon]. \quad (2.7)$$

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The dominance of the $1/(p \cdot q)$ pole in (2.5) implies that the effects of attaching one soft-graviton line to an arbitrary diagram is to supply a factor equal to the sum of (2.5) over all external lines in the diagram

$$(8\pi G)^{1/2} \sum_n \eta_n p_n^\mu / [p_n \cdot q - i\epsilon]. \quad (2.7)$$

Excerpt from Weinberg 1965 containing the basics of Weinberg's soft graviton theorem, which establishes how the S-matrix element for an amplitude with an extra soft graviton relates to the original S-matrix element.

In either case, so long as the source is not at a cosmologically large distance, the permanent change in the gravitational-wave field (the burst's memory) δh_{ij}^{TT} is equal to the 'transverse, traceless (TT) part' of time-independent, Coulomb-type, $1/r$ field of the final system minus that of the initial system. If \mathbf{P}^A is the 4-momentum of mass A of the system and P_i^A is a spatial component of that 4-momentum in the rest frame of the distant observer, and if \mathbf{k} is the past-directed null 4-vector from observer to source, then δh_{ij}^{TT} has the following form:

$$\delta h_{ij}^{TT} = \delta \left(\sum_A \frac{4P_i^A P_j^A}{\mathbf{k} \cdot \mathbf{P}^A} \right)^{TT} \quad (1)$$

Here we use units with $G = c = 1$.

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In either case, so long as the source is not at a cosmologically large distance, the permanent change in the gravitational-wave field (the burst's memory) δh_{ij}^{TT} is equal to the transverse, traceless (TT) part of time-independent, Coulomb-type $1/r$ field of the final system minus that of the initial system. If \mathbf{P}^A is the 4-momentum of mass A of the system and P^A is a spatial component of that 4-momentum in the rest frame of the distant observer, and if \mathbf{k} is the past-directed null 4-vector from observer to source, then δh_{ij}^{TT} has the following form:

$$\delta h_{ij}^{TT} = 6 \left(\sum_A \frac{4P^A P^A}{\mathbf{k} \cdot \mathbf{p}^A} \right)^{TT} \quad (1)$$

Here we use units with $G = c = 1$.

Excerpt from Braginsky and Thorne 1987 with the original form of the Braginsky-Thorne formula.

The Braginsky–Thorne formula is
a Fourier transform of Weinberg’s soft factor!

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The Braginsky–Thorne formula is
a Fourier transform of Weinberg's soft factor!

It was pointed out by Strominger and Zhiboedov (2016) that the Braginsky–Thorne formula is related to Weinberg's soft graviton factor by a Fourier transform.

Is there a symmetry behind Weinberg's
soft graviton theorem?

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Is there a symmetry behind Weinberg's
soft graviton theorem?

Relations between S -matrix elements are often due to symmetries. Is there a symmetry behind Weinberg's soft graviton theorem?

Bondi–Metzner–Sachs: at infinity in asymptotically flat spacetimes, the group of symmetries is the Poincaré group plus infinitely many **supertranslations**

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Bondi–Metzner–Sachs: at infinity in asymptotically flat spacetimes, the group of symmetries is the Poincaré group plus infinitely many supertranslations

Bondi, Metzner, and Sachs found out in the 1960s that the group of symmetries at null infinity in asymptotically flat spacetimes is not only the Poincaré group, but also includes infinitely many supertranslations. The full infinite-dimensional group is known as the Bondi–Metzner–Sachs (BMS) group.

Weinberg's soft graviton theorem follows from BMS invariance!

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Weinberg's soft graviton theorem
follows from BMS invariance!

Strominger (2014) and He et al. (2015) showed that Weinberg's soft graviton theorem is actually the Ward identity of the supertranslations.

The memory effect is
a physical realization of a supertranslation!

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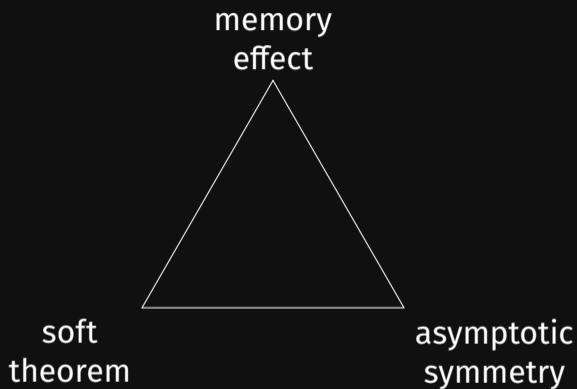
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The memory effect is
a physical realization of a supertranslation!

It was shown by Strominger and Zhiboedov (2016) that the displacement due to the memory effect can be understood as a supertranslation acting on the detectors.



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There is a relation between the memory effect, the soft graviton theorem, and the BMS symmetries. This triangle echoes through other theories and effects.

Memory and supertranslations may shed light on black hole evaporation

Strominger and Zhiboedov 2016. *J. High Energ. Phys.* **2016**, 86 . arXiv: 1411.5745 [hep-th]
Hawking, Perry, and Strominger 2016. *Phys. Rev. Lett.* **116**, 231301 . arXiv: 1601.00921 [hep-th]
Hawking, Perry, and Strominger 2017. *J. High Energ. Phys.* **2017**, 161 . arXiv: 1611.09175 [hep-th]

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Memory and supertranslations may shed light on black hole evaporation

Strominger and Zhiboedov (2016), Hawking, Perry, and Strominger (2016), and Hawking, Perry, and Strominger (2017) argued that supertranslations and the memory effect can shed some light on the loss of information through black hole evaporation.

A Few Comments on Measuring Memory





In binary mergers, the linear memory vanishes because the system is bound. However, the nonlinear effect contributes.




The memory is too faint for LIGO to detect it in individual events because it takes time for the memory to build up and the detectors are not inertial. Nevertheless, there are estimates that stacking many events together can lead to a detection in O5





LISA has a better chance at measuring memory in individual events because it responds better to low frequencies and it is inertial






Further Reading



- **Linear and nonlinear memory effects:** Bieri and A. Polnarev 2024. *Gravitational Wave Displacement and Velocity Memory Effects*. arXiv: 2402.02594 [gr-qc]
- **Relation with soft theorems and asymptotic symmetries:** Strominger 2018. *Lectures on the Infrared Structure of Gravity and Gauge Theory*. arXiv: 1703.05448 [hep-th]
- **Infrared divergences and soft theorems:** Weinberg 1995. *Foundations*

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